

# Principle Sources of Noise

- **Image Acquisition**
  - Image sensors may be affected by Environmental conditions (light levels etc)
  - Quality of Sensing Elements (can be affected by e.g. temperature)
- **Image Transmission**
  - Interference in the channel during transmission e.g. lightening and atmospheric disturbances

# Noise Model Assumptions

- Independent of Spatial Coordinates
- Uncorrelated with the image i.e. no correlation between Pixel Values and the Noise Component

# White Noise

- When the Fourier Spectrum of noise is constant the noise is called White Noise
- The terminology comes from the fact that the white light contains nearly all frequencies in the visible spectrum in equal proportions
- The Fourier Spectrum of a function containing all frequencies in equal proportions is a constant

# Noise Models: Gaussian Noise

Spatial noise descriptor based on the statistical behavior of the gray-level values  $\Rightarrow$  consider the gray-level values as random variables characterized by a probability density function (PDF)

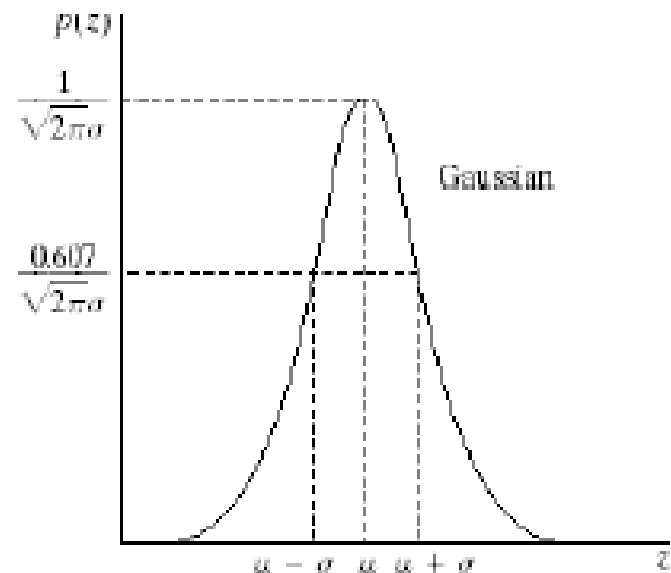
► Gaussian noise (also called “normal noise model”)

$$p(z) = \frac{1}{\sqrt{2\pi}\sigma} e^{-\frac{(z-\mu)^2}{2\sigma^2}}$$

$z$ : gray level

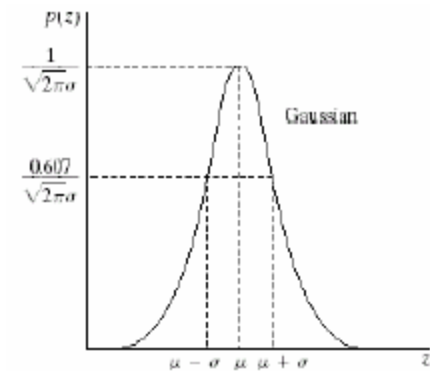
$\mu$ : mean of random variable  $z$

$\sigma^2$ : variance of  $z$



# Noise Models: Gaussian Noise

- Approximately 70% of its value will be in the range  $[(\mu-\sigma), (\mu+\sigma)]$  and about 95% within range  $[(\mu-2\sigma), (\mu+2\sigma)]$
- Gaussian Noise is used as approximation in cases such as Imaging Sensors operating at low light levels

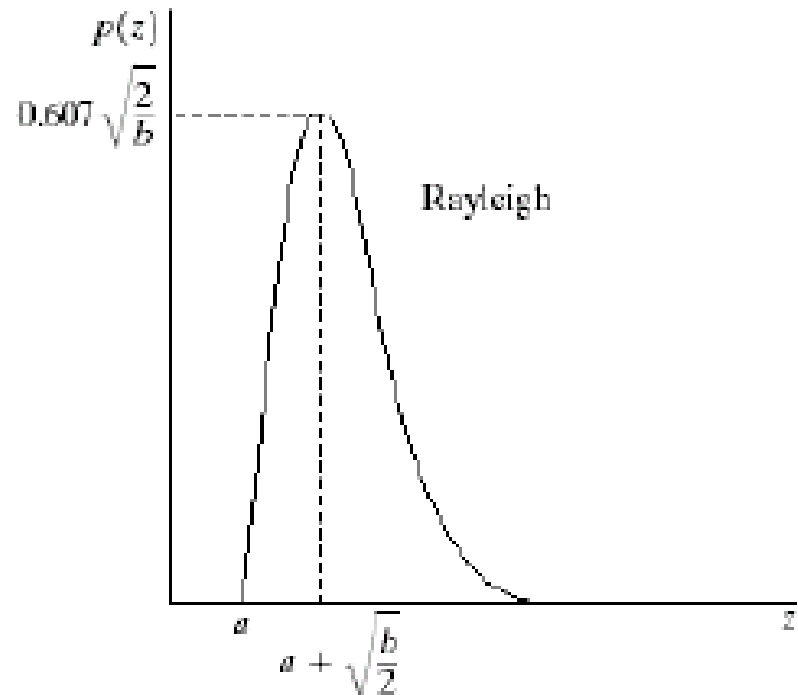


# Noise Models: Rayleigh Noise

$$p(z) = \begin{cases} \frac{2}{b}(z-a)e^{-\frac{(z-a)^2}{b}}, & \text{for } z \geq a \\ 0, & \text{for } z < a \end{cases}$$

$$\text{mean: } \mu = a + \sqrt{\frac{\pi b}{4}}$$

$$\text{variance: } \sigma^2 = \frac{b(4-\pi)}{4}$$



Rayleigh density can be used to approximate skewed histograms

Rayleigh Noise arises in Range Imaging

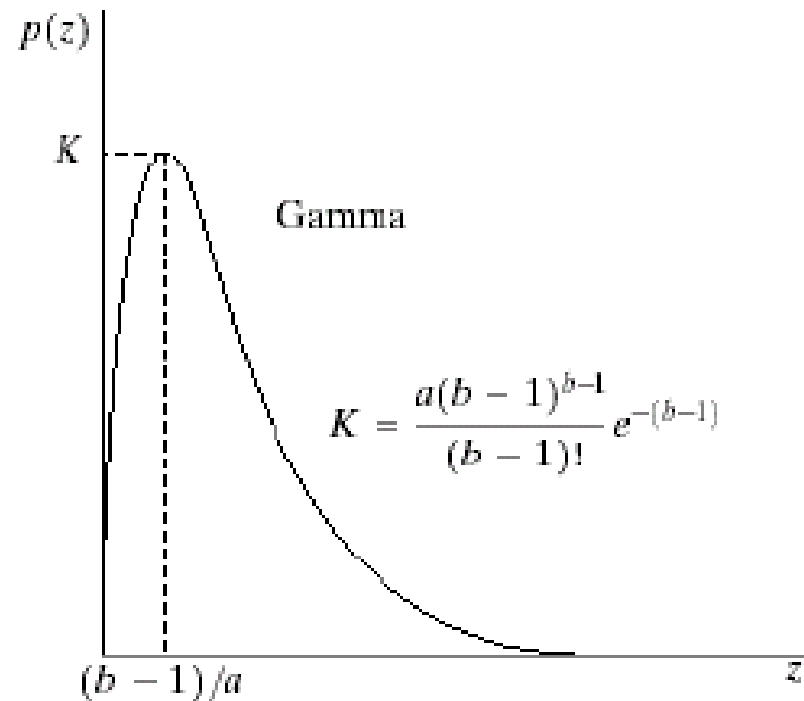
# Noise Models: Erlang (Gamma) Noise

$$p(z) = \begin{cases} \frac{a^b z^{b-1}}{(b-1)!} e^{-az}, & \text{for } z \geq 0 \\ 0, & \text{for } z < 0 \end{cases}$$

$$a > 0, b \in \mathbb{I}^+$$

$$\text{mean: } \mu = \frac{b}{a}$$

$$\text{variance: } \sigma^2 = \frac{b}{a^2}$$



Rayleigh Noise arises in Laser Imaging

# Noise Models: Exponential Noise

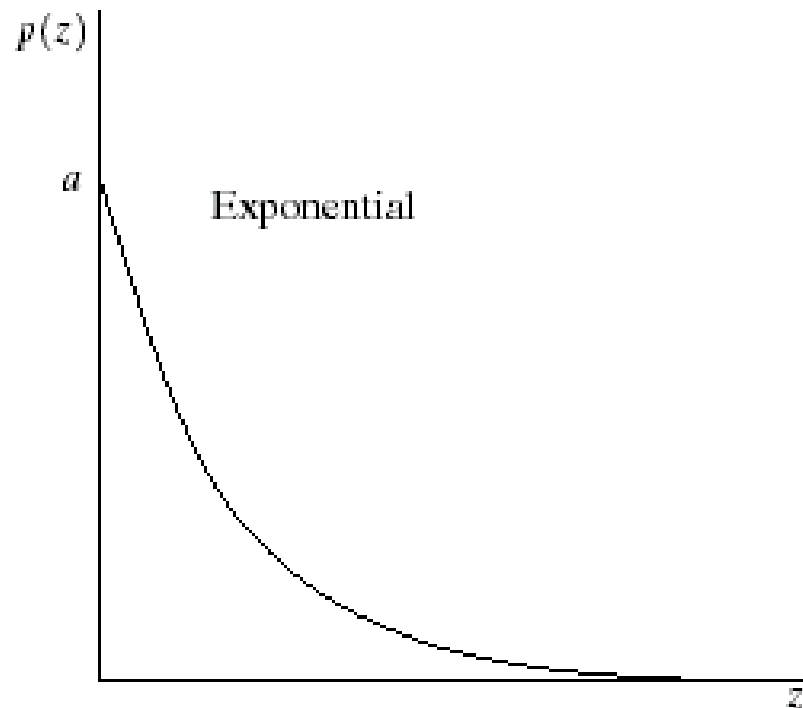
Special case of the Erlang PDF ( $b=1$ )

$$p(z) = \begin{cases} ae^{-az}, & \text{for } z \geq 0 \\ 0, & \text{for } z < 0 \end{cases}$$

Where  $a > 0$ ,

$$\text{mean: } \mu = \frac{1}{a}$$

$$\text{variance: } \sigma^2 = \frac{1}{a^2}$$



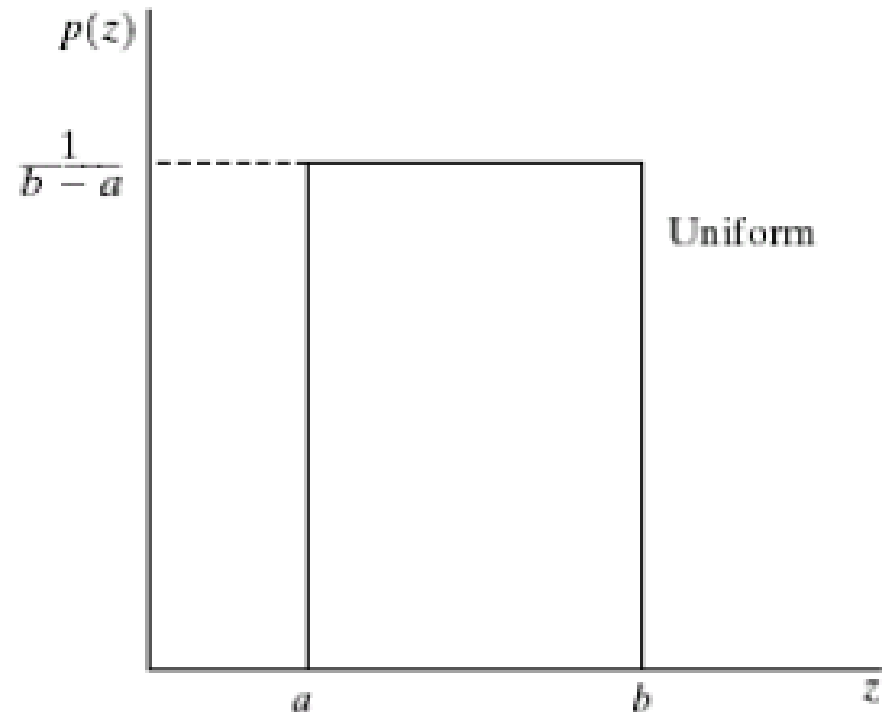


# Noise Models: Uniform Noise

$$p(z) = \begin{cases} \frac{1}{b-a}, & \text{if } a \leq z \leq b \\ 0 & \text{otherwise} \end{cases}$$

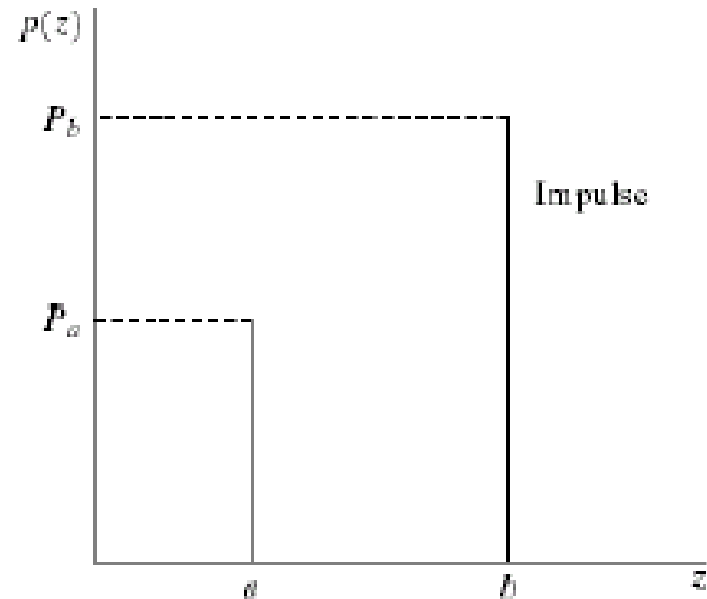
The mean and variance are given by

$$\mu = \frac{a+b}{2}, \quad \sigma^2 = \frac{(b-a)^2}{12}$$



# Noise Models: Impulse (Salt and Pepper) Noise

$$p(z) = \begin{cases} P_a & \text{for } z = a \\ P_b & \text{for } z = b \\ 0 & \text{otherwise} \end{cases}$$



If either  $P_a=0$  or  $P_b=0 \Rightarrow$  unipolar impulse noise

If  $P_a \approx P_b \neq 0 \Rightarrow$  bipolar impulse noise or salt-and-pepper noise  $\Rightarrow$  normally,  $a=0$  (black) and  $b=255$  (white)

# Applicability of Various Noise Models

Gaussian noise  $\Rightarrow$  electronic circuit noise and sensor noise due to poor illumination and/or high temperature

Rayleigh density  $\Rightarrow$  characterize noise phenomena in range imaging

Exponential and gamma densities  $\Rightarrow$  laser imaging

Impulse noise  $\Rightarrow$  occur when quick transients (faulty switching) take place during imaging

Uniform density  $\Rightarrow$  the least descriptive of practical situations

# Noise Models

